

Modelling turbulent fluxes due to thermal convection in rectilinear shearing flow

Radoslaw Smolec¹, Günter Houdek¹ and Douglas Gough²

¹Institute of Astronomy, University of Vienna, A-1180 Vienna, Austria
 email: radek.smolec@univie.ac.at, guenter.houdek@univie.ac.at

²Institute of Astronomy and Department of Applied Mathematics and Theoretical Physics,
 University of Cambridge, Cambridge CB3 0HA, UK
 email: douglas@ast.cam.ac.uk

Abstract. We revisit a phenomenological description of turbulent thermal convection along the lines proposed originally by Gough (1965) in which eddies grow solely by extracting energy from the unstably stratified mean state and are subsequently destroyed by internal shear instability. This work is part of an ongoing investigation for finding a procedure to calculate the turbulent fluxes of heat and momentum in the presence of a shearing background flow in stars.

Keywords. convection, turbulence, hydrodynamics

Introduction

Convection models based on the mixing-length approach still represent the main method for computing the turbulent fluxes in stars with convectively unstable regions. In such regions the pulsational stability of the star is affected not only by the radiative heat flux but also by the modulation of the convective heat flux and by direct mechanical coupling of the pulsation with the convective motion via the Reynolds stresses. Time-dependent formulations of the mixing-length approach for radial pulsation have been proposed by, for example, Gough (1965, 1977a) and Unno (1967). In a first step towards a generalization to nonradially pulsating stars, Gough & Houdek (2001) adopted Gough's formulation, incorporating into it a treatment of the influence of a shearing background flow. In this generalized framework of the mixing-length formalism, in which turbulent convective eddies grow according to linearized theory and are subsequently broken up by internal shear instability, there is a consequent reduction in the mean amplitude of the eddy motion, and a corresponding reduction in the heat flux.

In order to test and calibrate the formalism, it is preferable first to compare its predictions with existing results of hopefully more reliable investigations, such as experiments or numerical simulations. Here we extend our earlier work (Gough & Houdek 2001), comparing the functional forms of our mean temperature and shear profiles with those of direct numerical simulations (DNS; Domaradzki & Metcalfe 1988, DM88) of Rayleigh-Bénard convection in air (Prandtl number, $\sigma = 0.71$) with a strongly shearing background flow. We consider a plane-parallel layer of fluid confined between rigid horizontal perfectly conducting boundaries of infinite extent at fixed temperatures, the lower being hotter than the upper by ΔT . The upper boundary moves horizontally with constant velocity ΔU , and we assume, in accordance with the Boussinesq approximation, that the shear, E , in the mean (plane Couette) flow does not vary over the scale of an eddy.

Turbulent fluxes and mean equations in the presence of a shear

We follow the basic procedure by Gough & Houdek (2001) to solve the linearized equations describing the dynamics in a statistically stationary flow of a viscous Boussinesq fluid confined between two horizontal planes. In Cartesian co-ordinates (x, y, z) the equa-

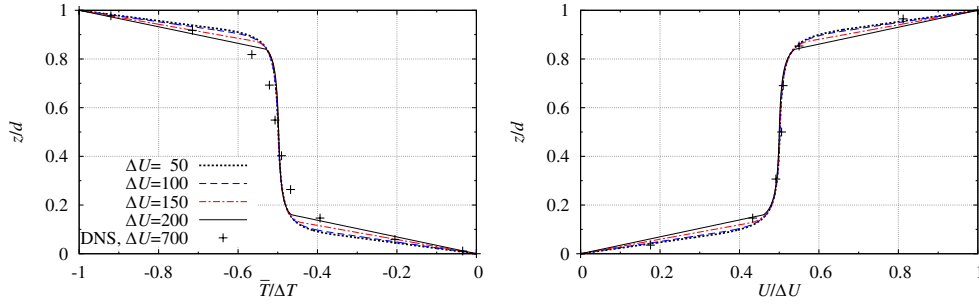


Figure 1. Mean temperature $\bar{T}/\Delta T$ (left panel) and mean velocity $U/\Delta U$ (right panel) as a function of height z are compared with DNS data (crosses) for different ΔU values.

tions are nondimensionalized using the vertical distance between the planes, d , and the thermal diffusion time across d as units of space and time. Linearized modes of convection are obtained by expansion about $E := d_z U = 0$ (where $d_z \equiv d/dz$) to second order in E . The resulting expressions for the eigenfunctions of the fluctuating temperature, T' , and turbulent velocity field, $\mathbf{u} = (u, v, w)$, are then used to compute the turbulent fluxes of heat, $\overline{wT'}$, and momentum (Reynolds stresses), $\overline{\rho u_i u_j}$, in the manner of Gough (1977a). A horizontal bar denotes statistical average (average over the horizontal plane). In our dimensionless formulation, the sum of the radiative and convective fluxes is independent of z and equal to the Nusselt number of the layer, $N = F_r(z) + F_c(z)$. In the diffusion approximation the radiative flux is equal to the mean temperature gradient, $F_r(z) = -d_z \bar{T}$. The thermal mean equation is solved simultaneously with (the only non-trivial (x -) component of) the mean momentum equation: $d_z \overline{\rho u w} - \sigma d_{zz} U = 0$ (where $d_{zz} \equiv d^2/dz^2$). Note that the Reynolds stress $\overline{\rho u w}$ distorts the shear, and consequently the x -component of the mean flow, U , is not a linear function of height z .

Results and conclusions

Fig. 1 shows the normalized mean vertical velocity profiles, $U/\Delta U$ and mean temperature profiles, $\bar{T}/\Delta T$, for four values of ΔU : 50, 100, 150 and 200. Our results are compared with DNS data (DM88), which assume $\Delta U = 700$. The mean profiles are in reasonable agreement with the DNS data, but for smaller values of ΔU in our model computations. Best agreement with the DNS data is obtained with $\Delta U = 150$. Two factors may be responsible for the discrepancy in the values of ΔU : (i) to maintain greatest simplicity, we adopted horizontal-stress-free boundary conditions for the eddies, whereas in the simulations the bounding planes were taken to be rigid – this may perhaps account for up to a factor of about three in the ΔU differences between the model ($\Delta U=150$) and the simulation ($\Delta U=700$); (ii) nonlocal effects may also contribute to the remaining differences between our results and the DNS data. We plan to investigate these issues, extending our model in the manner of Gough (1977b) to accommodate nonlocal behaviour.

This research is supported by the Austrian FWF grant No. AP 2120521. DOG is grateful to the Leverhulme Foundation for an Emeritus Fellowship.

References

- Domaradzki, J.A. & Metcalfe, R.W. 1988, *J. Fluid Mech.*, 193, 499
- Gough, D.O. 1965, *Geophys. Fluid Dyn. II*, Woods Hole Oceanographic Institution, p.49
- Gough, D.O. 1977a, *ApJ*, 214, 196
- Gough, D.O. 1977b, *Problems of stellar convection*, Springer-Verlag, Berlin, p. 15
- Gough, D.O. & Houdek, G. 2001, *ESASP*, 464, 637
- Unno W. 1967, *PASJ*, 19, 140